

# Proton scalar polarizabilities at MAMI

Frontiers and Careers in Photonuclear Physics 2019

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Edoardo Mornacchi

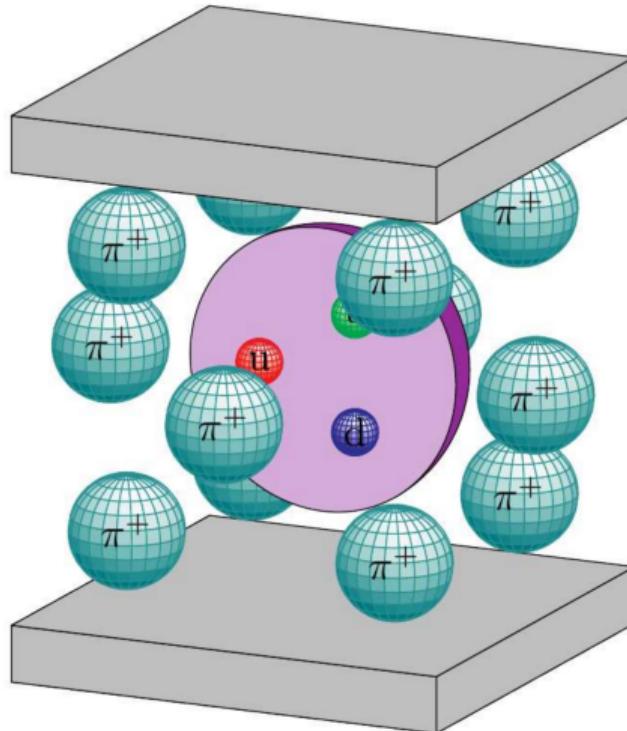
on behalf of the A2 collaboration

Pafos, October 28<sup>th</sup> 2019

Institute for Nuclear Physics  
Johannes Gutenberg University of Mainz



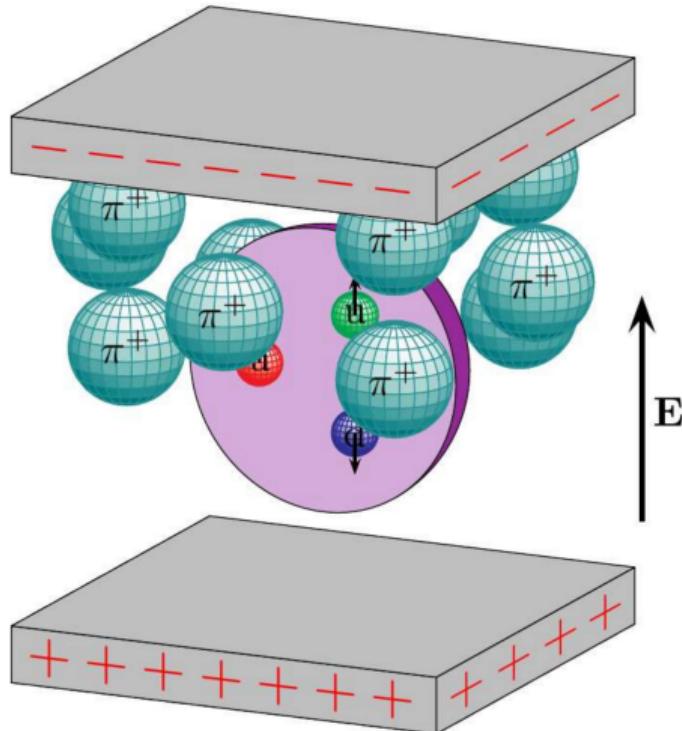
# Electric scalar polarizability - $\alpha_{E1}$



Describes the response of a proton to an applied electric field:

Picture: P. Martel

# Electric scalar polarizability - $\alpha_{E1}$



Describes the response of a proton to an applied electric field:

- Electric dipole moment:

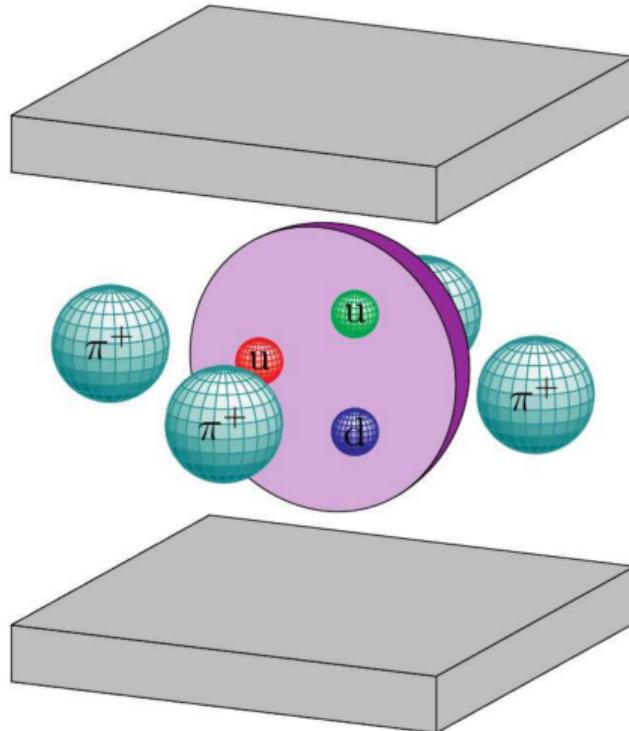
$$\vec{p} = [\alpha_{E1}] \times \vec{E}$$

Electric polarizability

- “Stretchability” of the proton

Picture: P. Martel

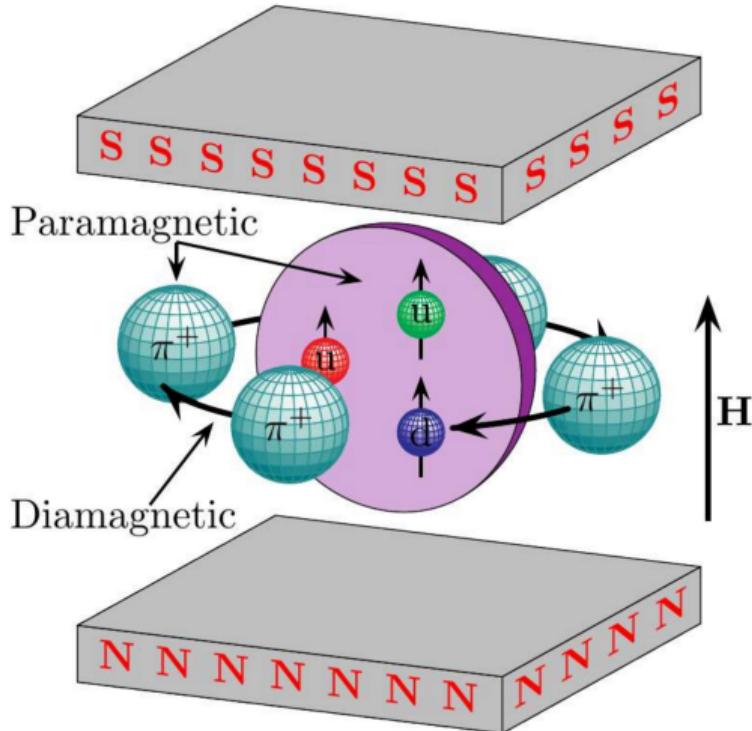
# Magnetic scalar polarizability - $\beta_{M1}$



Describes the response of a proton to an applied magnetic field:

Picture: P. Martel

# Magnetic scalar polarizability - $\beta_{M1}$



Picture: P. Martel

Describes the response of a proton to an applied magnetic field:

- Magnetic dipole moment:

$$\vec{m} = \boxed{\beta_{M1}} \times \vec{H}$$

Magnetic polarizability

- “Alignability” of the proton

# Measure proton scalar polarizabilities



Why measure them?

- Fundamental properties related to nucleon internal structure
- Limit precision to different area of physics:
  - two-photon exchange contribution to the Lamb shift and hyperfine structure in atomic physics
  - determination of the EM contribution to n-p mass difference
  - neutron star susceptibility
- Fertile meeting ground between theory and experiment

# Measure proton scalar polarizabilities

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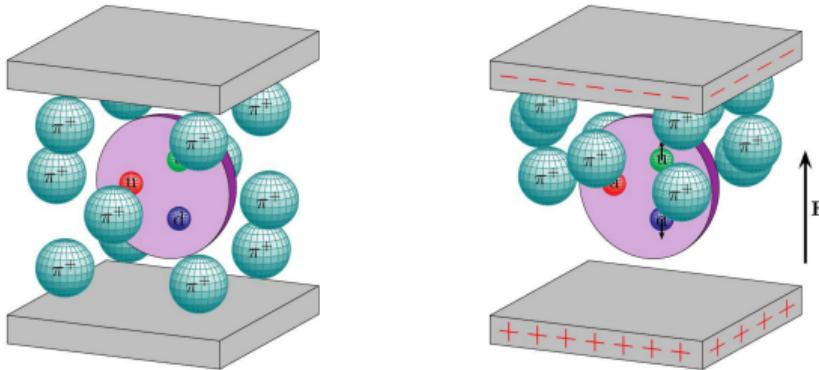
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- Fertile meeting ground between theory and experiment

OK!! But how?

# Measure proton scalar polarizabilities

How to measure them?

- We could place protons in a static electric field!



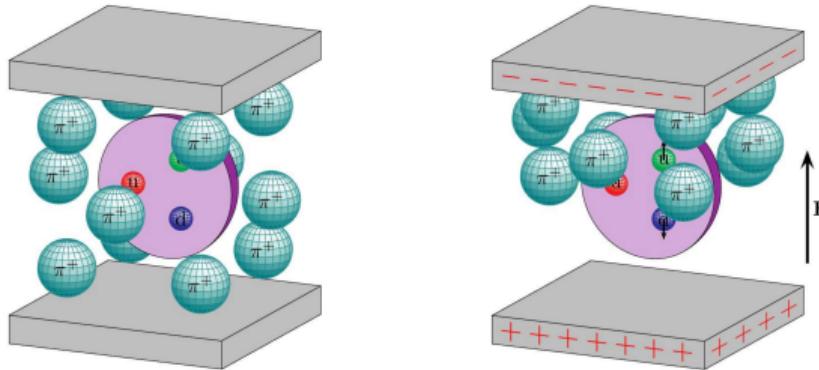
Picture: P. Martel

Ok, but how strong should it be to induce any appreciable polarizability?

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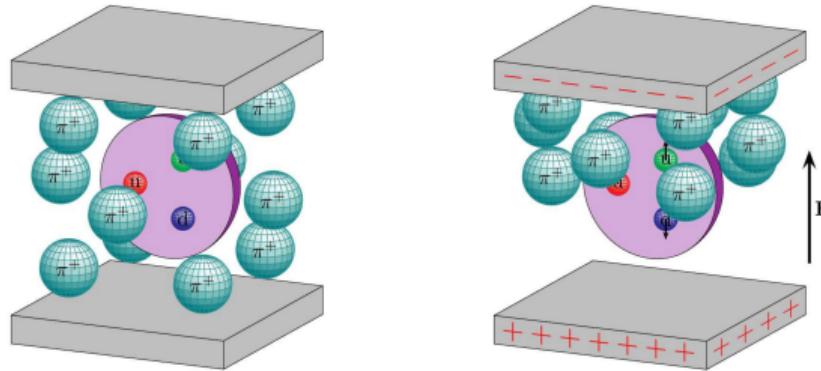
For an atom the polarizability effect is proportional to the volume:

$$\alpha_{E_1} \propto r^3$$

# Measure proton scalar polarizabilities

How to measure them?

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Picture: P. Martel

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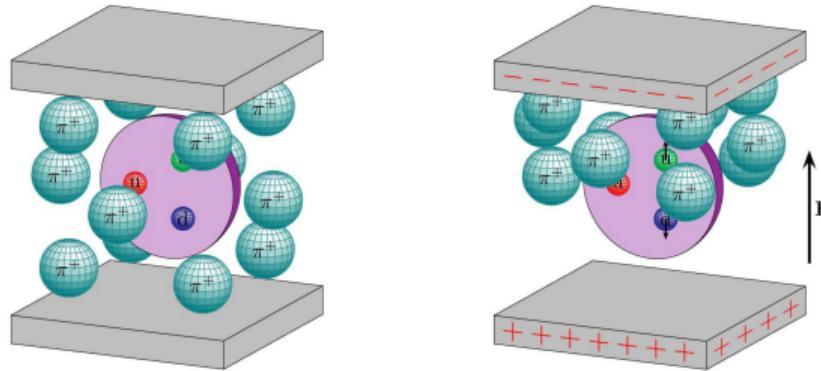
Assuming  $r_p \approx 0.875$  fm, the polarizability should be:

$$\alpha_{E_1} \propto r^3 \approx 0.6 \text{ fm}^3$$

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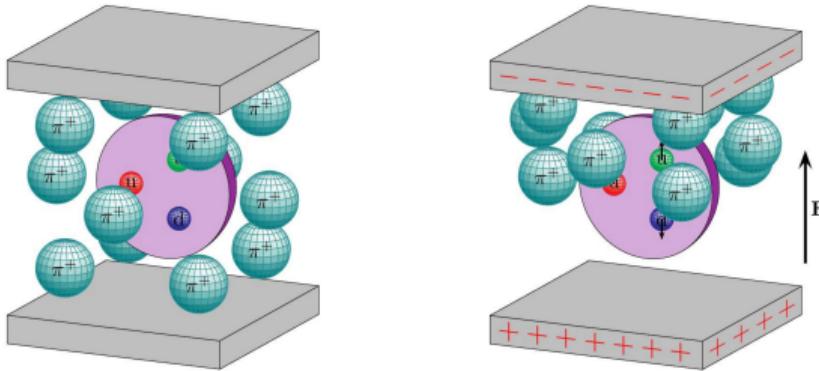
Assuming  $r_p \approx 0.875$  fm, the polarizability should be:

$$\alpha_{E_1} \propto r^3 \approx 0.6 \text{ fm}^3, \text{ but } \alpha_{E_1}^{\text{exp}} \approx 10 \times 10^{-4} \text{ fm}^3$$

# Measure proton scalar polarizabilities

How to measure them?

- We could place protons in a static electric field!



Picture: P. Martel

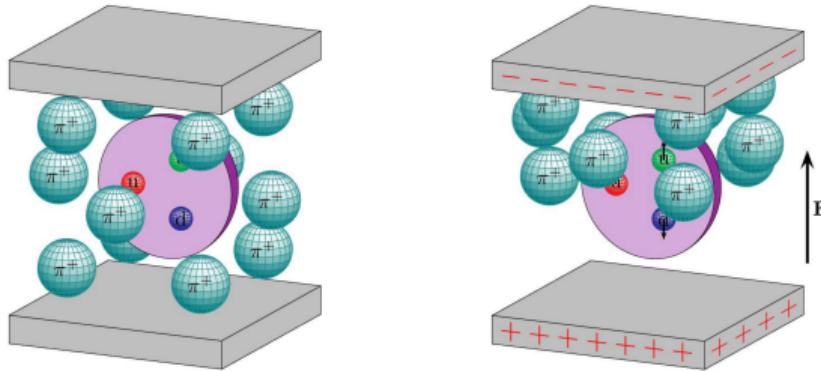
Ok, but how strong should it be to induce any appreciable polarizability?

Nucleon mass is mostly coming from binding force!  
⇒ It is really stiff and strongly bound!

# Measure proton scalar polarizabilities

How to measure them?

- We could place protons in a static electric field!



Picture: P. Martel

Ok, but how strong should it be to induce any appreciable polarizability?

From the energy level spacing and size of the nucleon, one can estimate:  $E_{\text{crit}} \approx 10^{23} \text{ V/m}$

# Measure proton scalar polarizabilities

How to measure them?

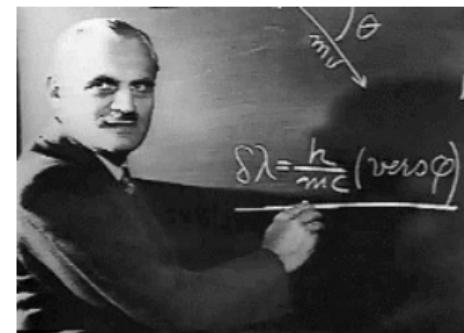
- ✗ We could **NOT** place protons in a static electric field!

# Measure proton scalar polarizabilities

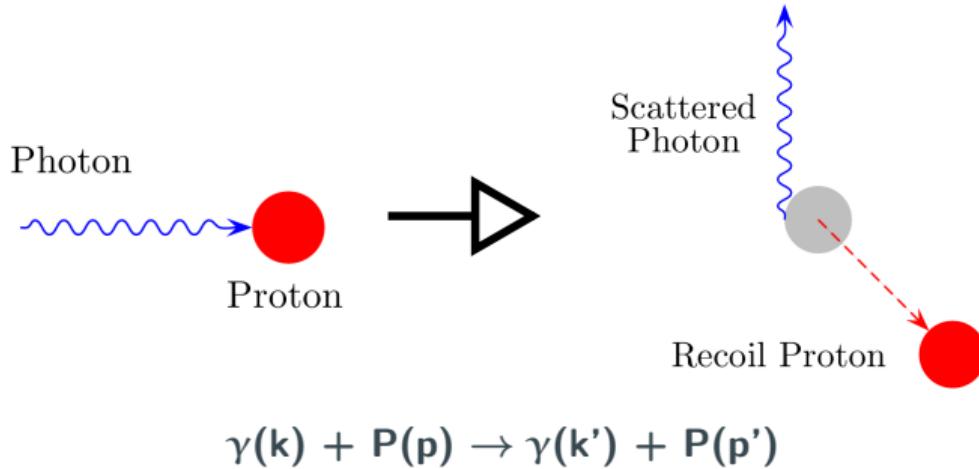
How to measure them?

✗ We could **NOT** place protons in a static electric field!

- We can use **Compton scattering!**
  - Observed for the first time by Arthur Compton in 1923
  - No classical explanation
  - Clear evidence of relativity and particle-like nature of the light



# Nuclear Compton scattering



Internal structure of the proton can be accessed by measuring unpolarized cross-section and polarization observables for Compton scattering

## Born term

Under the assumption of NO proton internal structure, the effective Hamiltonian can be written in terms of mass, electric charge and anomalous magnetic moment

- Zeroth order: mass and electric charge

$$H_{\text{eff}}^{(0)} = \frac{\vec{\pi}^2}{2m} + e\phi \quad (\text{where } \vec{\pi} = \vec{p} - e\vec{A})$$

- First order: anomalous magnetic moment

$$H_{\text{eff}}^{(1)} = -\frac{e(1+k)}{2m}\vec{\sigma} \cdot \vec{H} - \frac{e(1+2k)}{8m^2}\vec{\sigma} \cdot [\vec{E} \times \vec{\pi} - \vec{\pi} \times \vec{E}]$$

## Scalar polarizabilities

Effective Hamiltonian at the second order includes scalar polarizabilities, which are related to the proton internal structure

- Second order: scalar polarizabilities  $\alpha_{E1}$  and  $\beta_{M1}$

$$H_{\text{eff}}^{(2)} = -4\pi \left[ \frac{1}{2} \alpha_{E1} \vec{E}^2 + \frac{1}{2} \beta_{M1} \vec{H}^2 \right]$$

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Interested in what does happen for higher orders?

**P. Martel** talk: Wed., Workshop 1, Session 1

## Scalar polarizabilities

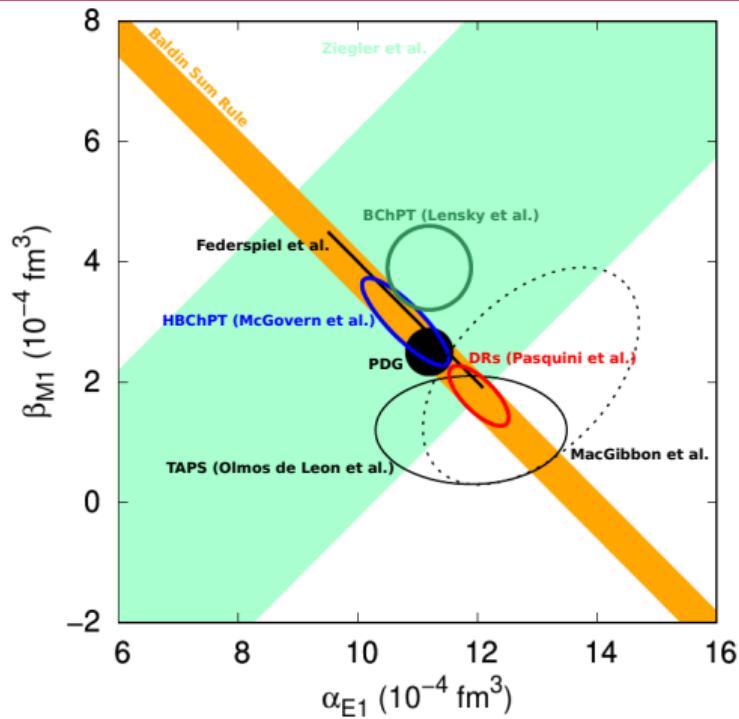
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$$H_{\text{eff}}^{(2)} = -4\pi \left[ \frac{1}{2} \alpha_{E1} \vec{E}^2 + \frac{1}{2} \beta_{M1} \vec{H}^2 \right]$$

How is the current situation for the scalar polarizabilities measurements?

# Existing data on scalar polarizabilities



B. Pasquini, P. Pedroni and S. Sconfietti, J. Phys. G 46, no. 10, 104001 (2019).

**PDG (2012) values:**

$$\alpha_{E1} = (12.0 \pm 0.6) 10^{-4} \text{ fm}^3$$

$$\beta_{M1} = (1.9 \pm 0.5) 10^{-4} \text{ fm}^3$$

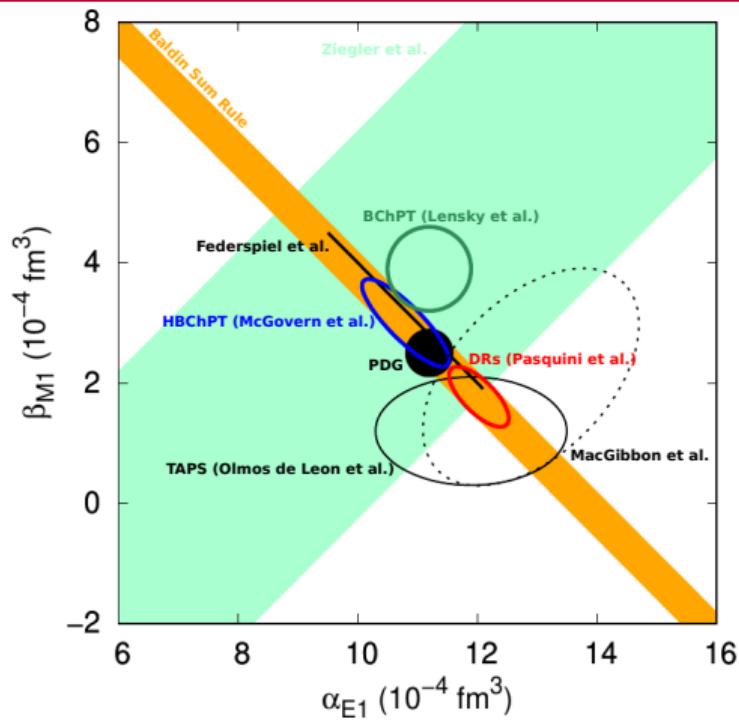
**Current PDG values:**

$$\alpha_{E1} = (11.2 \pm 0.4) 10^{-4} \text{ fm}^3$$

$$\beta_{M1} = (2.5 \pm 0.4) 10^{-4} \text{ fm}^3$$

Significant change between reviews without new experimental data  
 ⇒ Dataset not fully consistent!

# Existing data on scalar polarizabilities



B. Pasquini, P. Pedroni and S. Sconfietti, J. Phys. G 46, no. 10, 104001 (2019).

⇒ New high-precision dataset is needed!

**PDG (2012) values:**

$$\alpha_{E1} = (12.0 \pm 0.6) 10^{-4} \text{ fm}^3$$

$$\beta_{M1} = (1.9 \pm 0.5) 10^{-4} \text{ fm}^3$$

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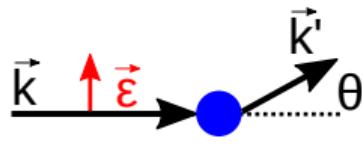
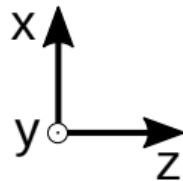
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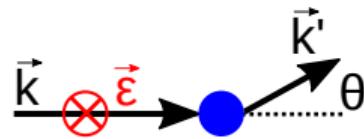
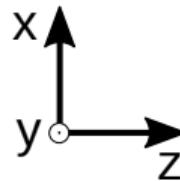
Significant change between reviews without new experimental data  
⇒ Dataset not fully consistent!

- World existing dataset was previously obtained using only unpolarized cross-section for Compton scattering
- At low energy, below the pion photoproduction threshold, the measurement of the beam asymmetry  $\Sigma_3$  provides an alternative way to extract  $\beta_{M1}$

$$\Sigma_3 = \frac{d\sigma_{||} - d\sigma_{\perp}}{d\sigma_{||} + d\sigma_{\perp}}$$

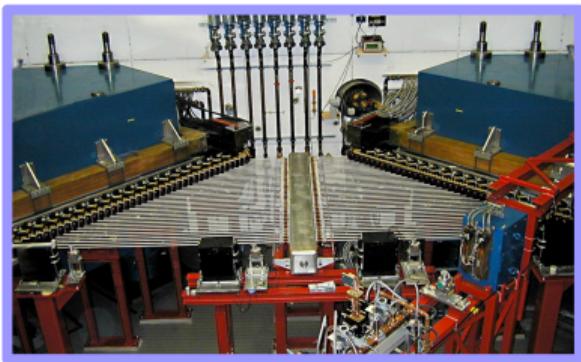
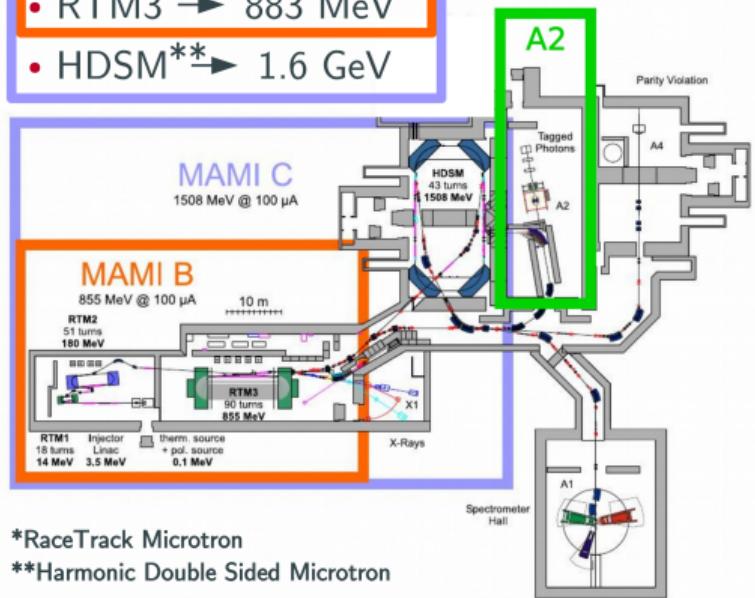


PARA(LLEL)



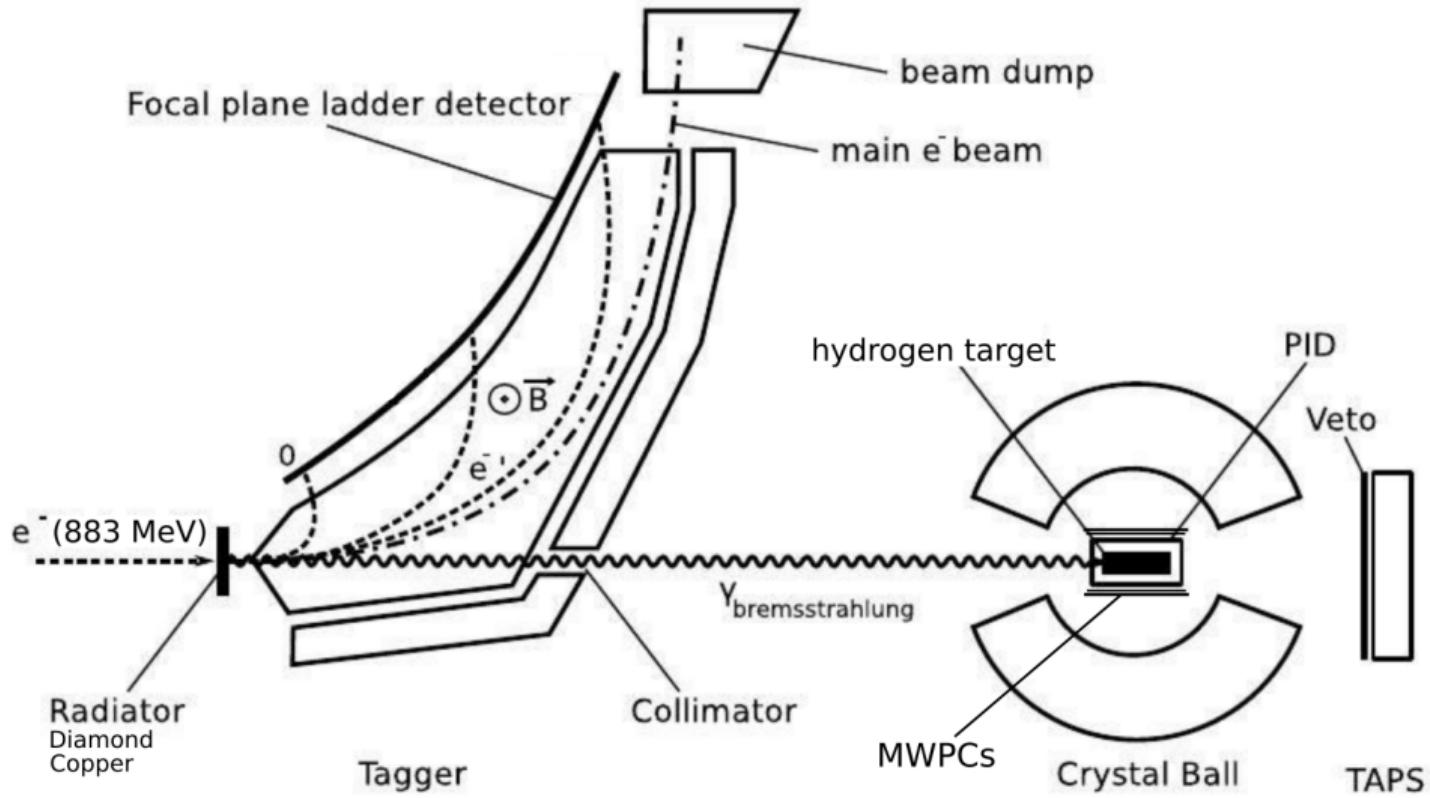
PERP(ENDICULAR)

- Injector → 3.5 MeV
- RTM1\* → 14.9 MeV
- RTM2 → 180 MeV
- RTM3 → 883 MeV
- HDSM\*\* → 1.6 GeV



\*RaceTrack Microtron

\*\*Harmonic Double Sided Microtron



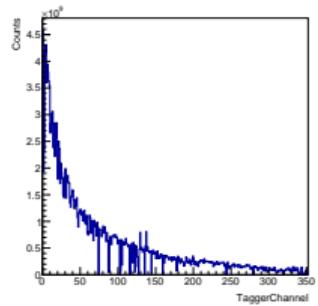
# Tagging system upgrade

A2

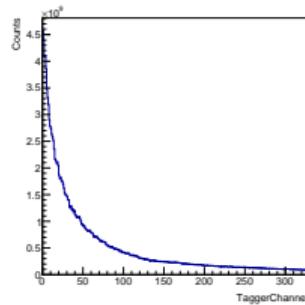
- Higher photon flux
- Higher efficiency
- Better control of systematic



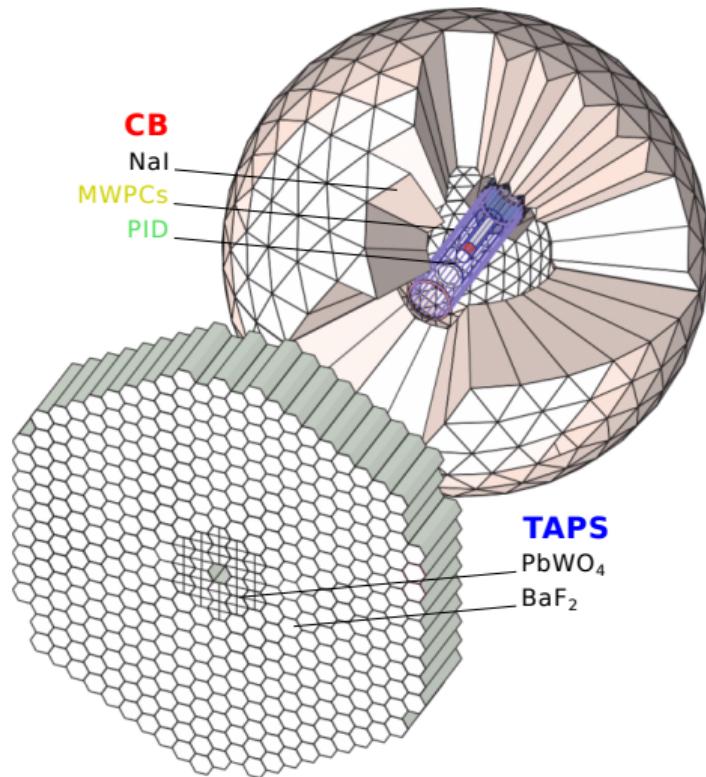
Hits distribution in the old tagger



Hits distribution in the new tagger



# Experimental setup



## Crystal Ball

- 672 NaI(Tl) crystals
- Particle Identification Detector (**PID**):  
24 scintillator paddles
- 2 Multiwire Proportional Chambers (**MWPCs**)

## TAPS

- 366 BaF<sub>2</sub> and  
72 PbWO<sub>4</sub> crystals
- 384 veto paddles

## Data collection:

- Pilot experiment: data collected in June 2013
- New high precision experiment: data collected in the first half of 2018

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We are selecting **Compton scattering**  $\vec{\gamma}p \rightarrow \gamma p$  events with:

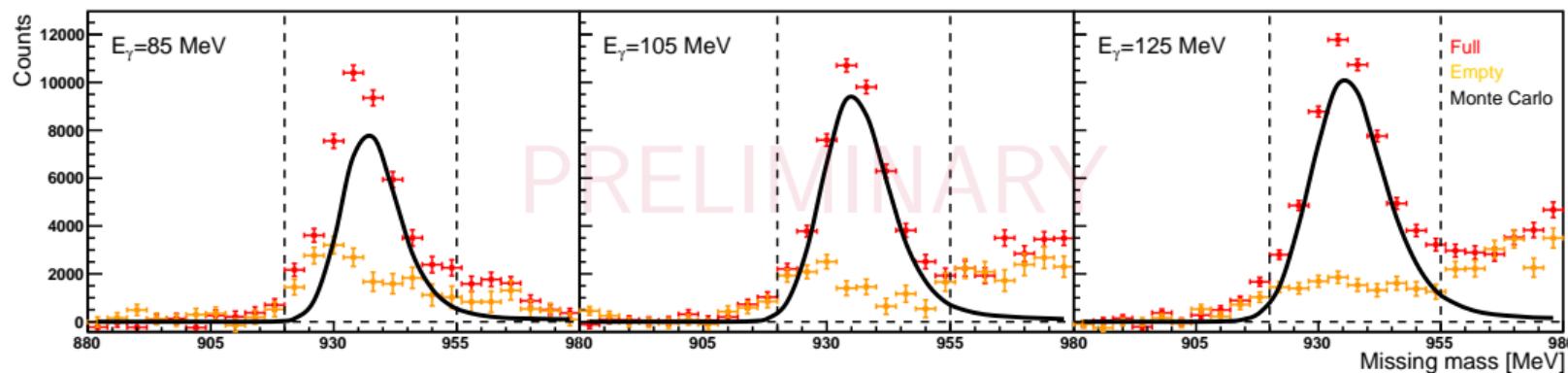
$$\Rightarrow E_{\gamma_{\text{beam}}} = 80 - 140 \text{ MeV} \text{ and } \theta_{\gamma_{\text{out}}} = 30^\circ - 155^\circ$$

- Subtraction of random coincidences in the tagger
- 1  $\gamma$  in the final state
- Subtraction of the empty target contribution
- Missing mass cut
- Linear polarization degree extraction event by event
- Constant flux monitoring using a pair spectrometer

# Missing mass distribution — PARALLEL dataset

A2

The LH<sub>2</sub> target requires separate data taking with the empty target to determine the contribution of the target cell itself

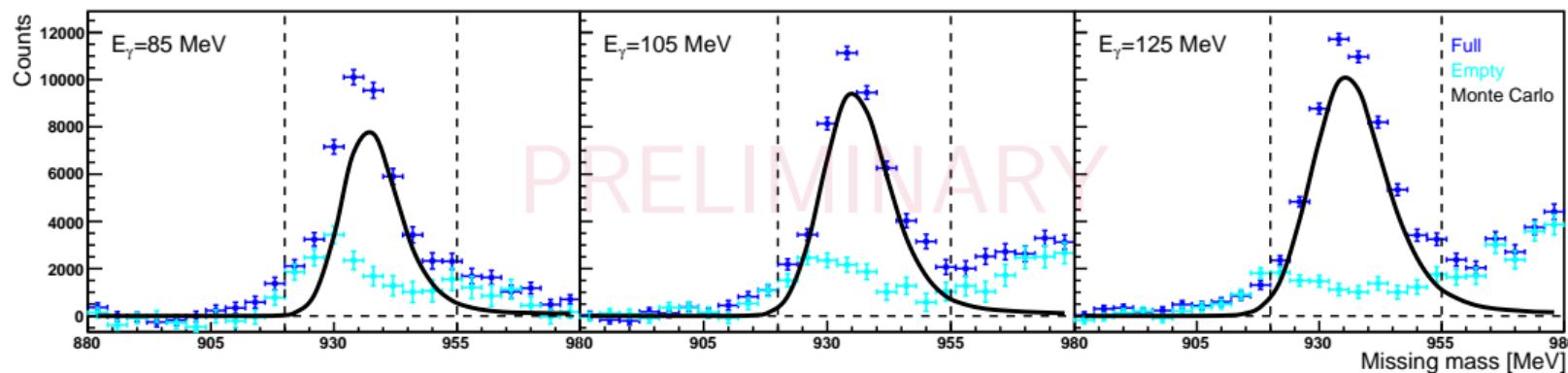


$$m_{\text{miss}} = \sqrt{(E_{\gamma_i} + m_p - E_{\gamma_f})^2 - (\vec{p}_{\gamma_i} - \vec{p}_{\gamma_f})^2} = m_p$$

# Missing mass distribution — PERPENDICULAR dataset

A2

The LH<sub>2</sub> target requires separate data taking with the empty target to determine the contribution of the target cell itself

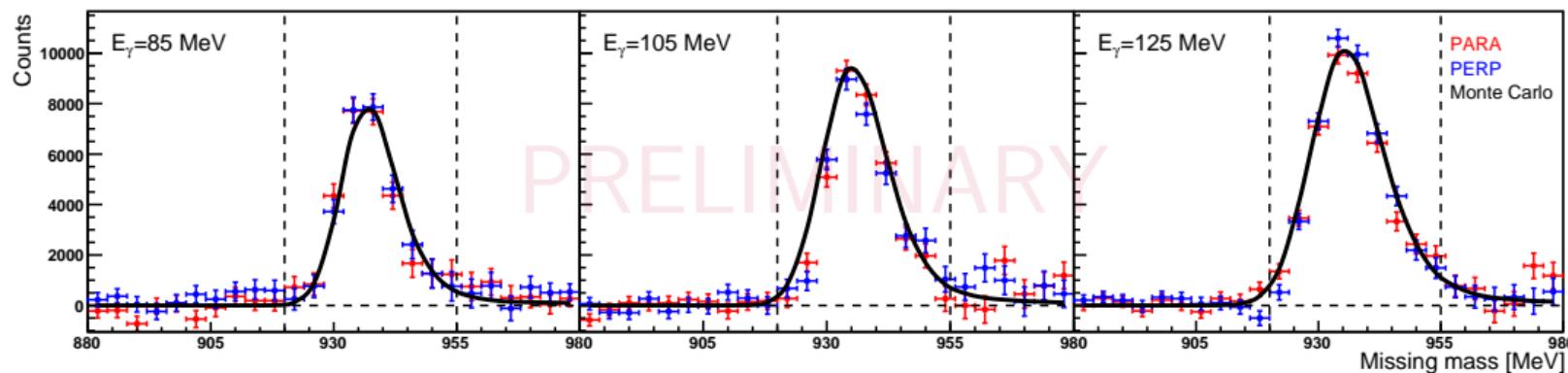


$$m_{\text{miss}} = \sqrt{(E_{\gamma_i} + m_p - E_{\gamma_f})^2 - (\vec{p}_{\gamma_i} - \vec{p}_{\gamma_f})^2} = m_p$$

# Missing mass distribution — after empty target subtraction

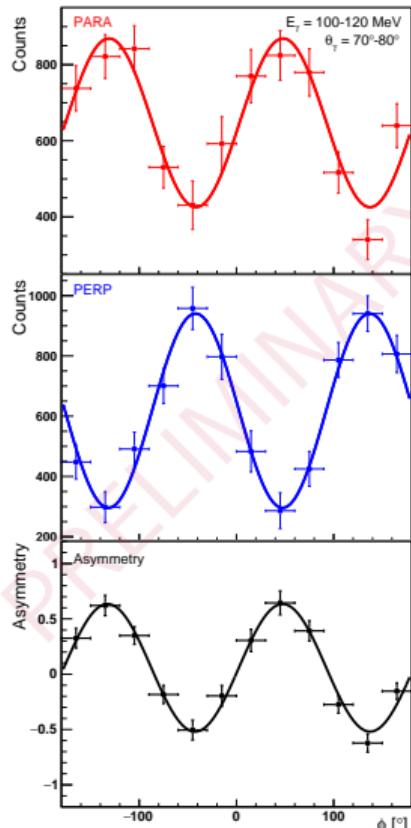
A2

Good agreement between PARA, PERP and Monte Carlo simulation.  
Very good statistics with low background!

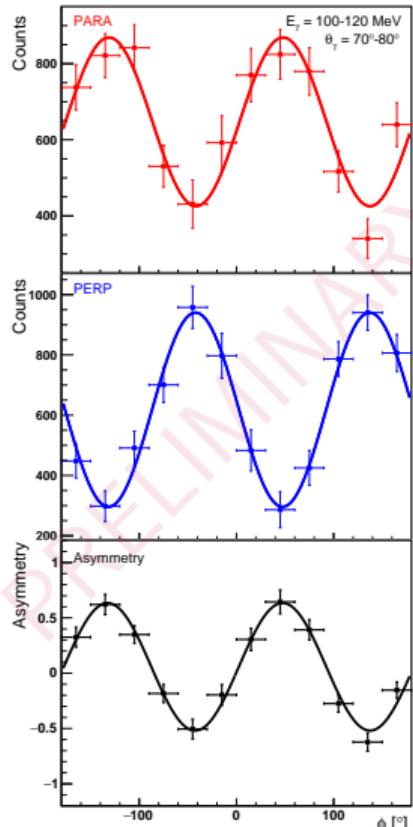


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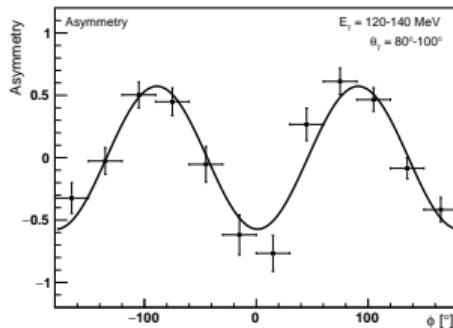
# New measurement - $\phi$ distributions



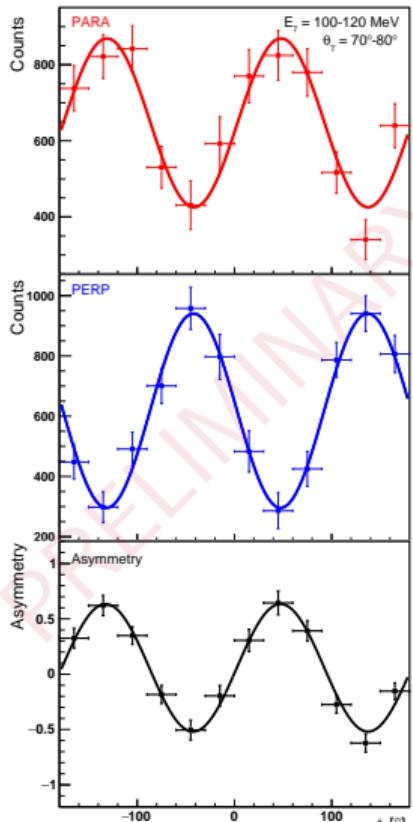
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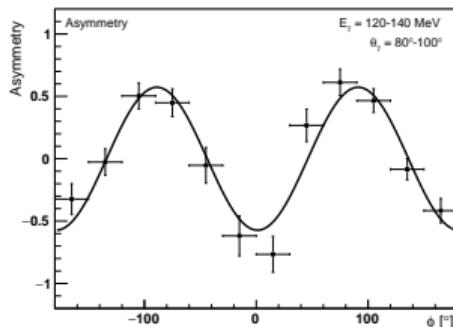
Example of  $\phi$  distribution from the pilot experiment dataset:



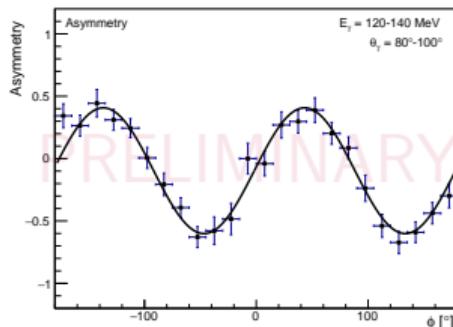
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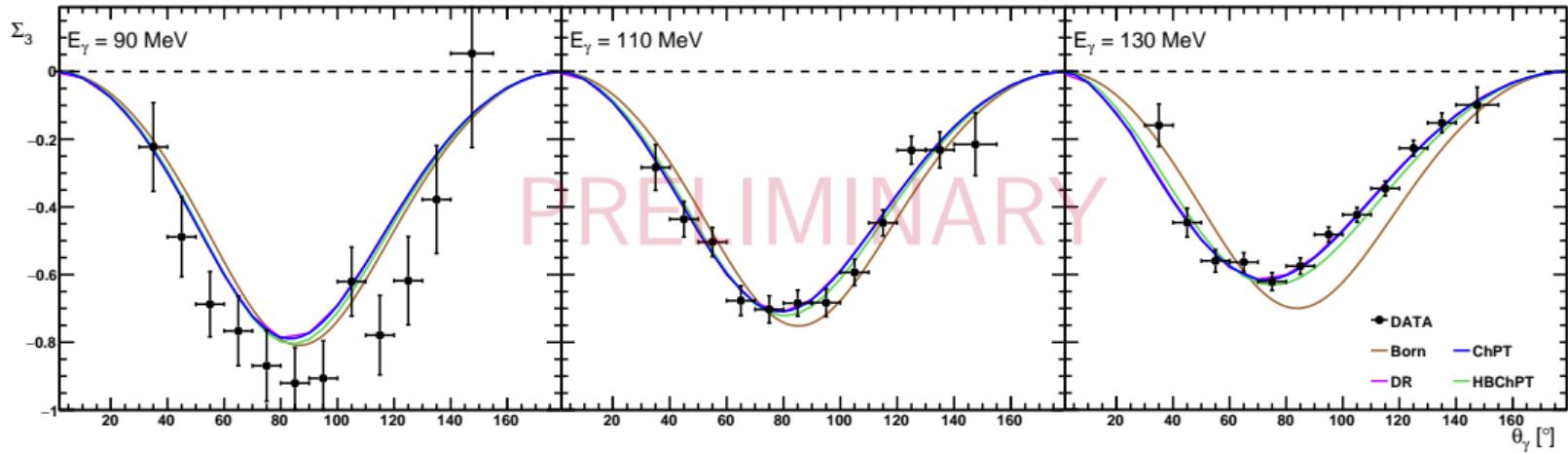
Example of  $\phi$  distribution from the pilot experiment dataset:



Same example for the new dataset, with double number of bins in  $\phi$ !!



# New measurement - Beam asymmetry $\Sigma_3$



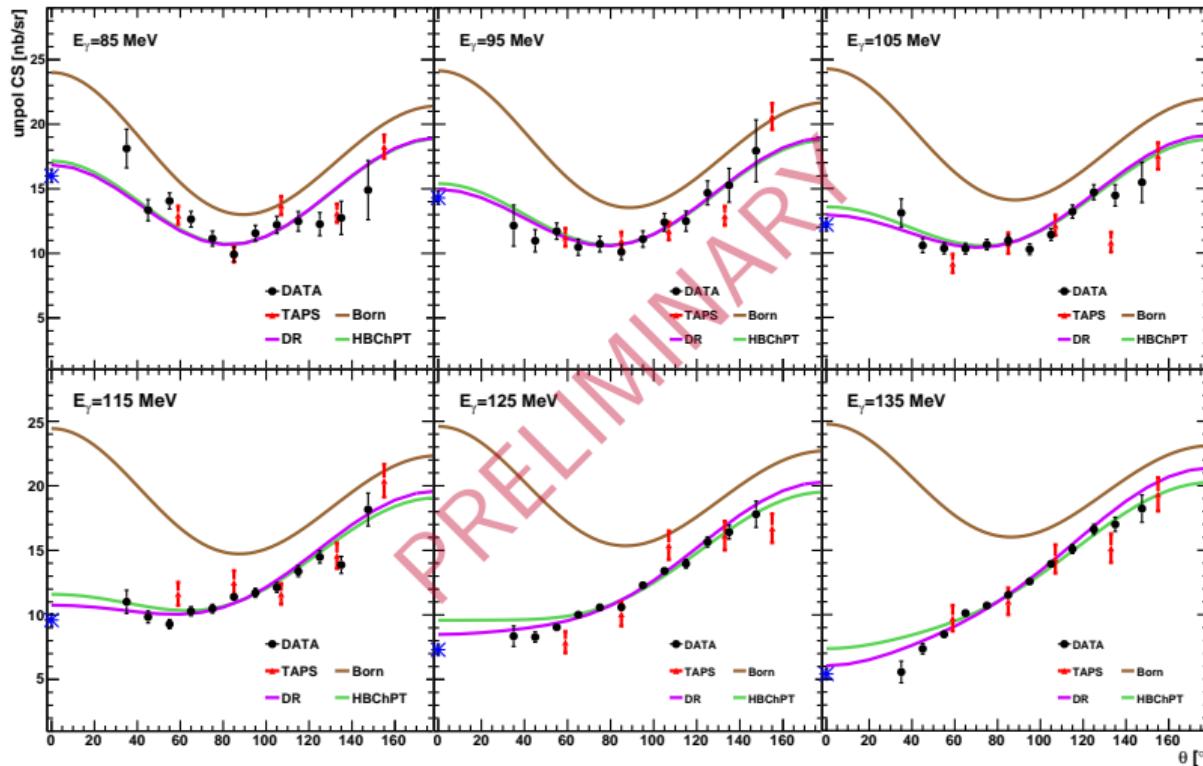
**Big improvement in statistics compared to the pilot experiment!**

Lensky, V. & Pascalutsa, V.,  
Eur. Phys. J. C (2010) 65:195

McGovern, J.A., Phillips, D.R. &  
Grießhammer, H.W.,  
Eur. Phys. J. A (2013) 49:12

B. Pasquini, D. Drechsel,  
& M. Vanderhaeghen,  
Phys. Rev. C 76

# New measurement - Unpolarized cross-section



Good agreement with theoretical predictions and improvement in statistics compared to TAPS dataset!

V. Olmos de Leon, et al.,  
*Eur. Phys. J. A* 10 (2001)  
 McGovern, J.A., Phillips, D.R. &  
 Grießhammer,  
 H.W., *Eur. Phys. J. A* (2013) 49: 12  
 B. Pasquini, D. Drechsel, and M.  
 Vanderhaeghen,  
*Phys. Rev. C* 76

- Successful first data taking with the new tagging system
- 1.2 million good Compton scattering events in the relevant energy range
- Simultaneous high precision measurement of unpolarized cross-section and  $\Sigma_3$
- Preliminary results are definitively very promising!
- Preliminary checks showed a small systematic error
- Preliminary fits for the extraction of the scalar polarizabilities showed a significant improvement compared to the biggest data-set currently published
- Analysis is almost finalized and a publication is expected soon

# THANKS!

A2

Special thanks to all the A2 collaboration members!



# THANKS!

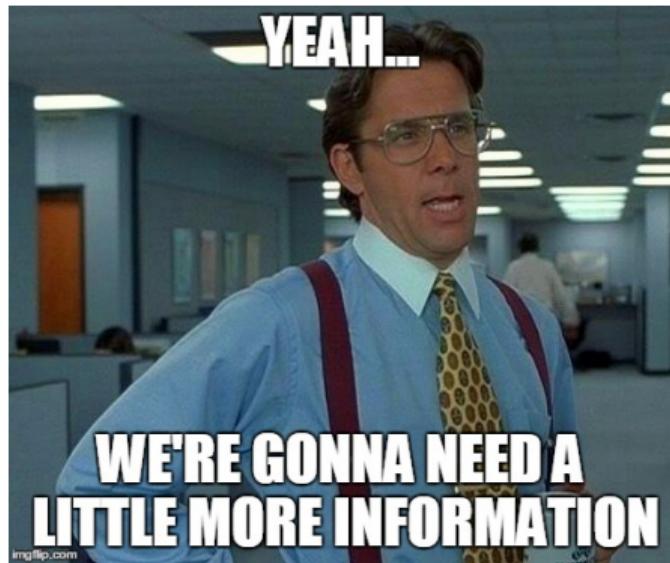
A2

Special thanks to all the A2 collaboration members!



and in particular...

**THANKS TO YOU FOR YOUR ATTENTION!**

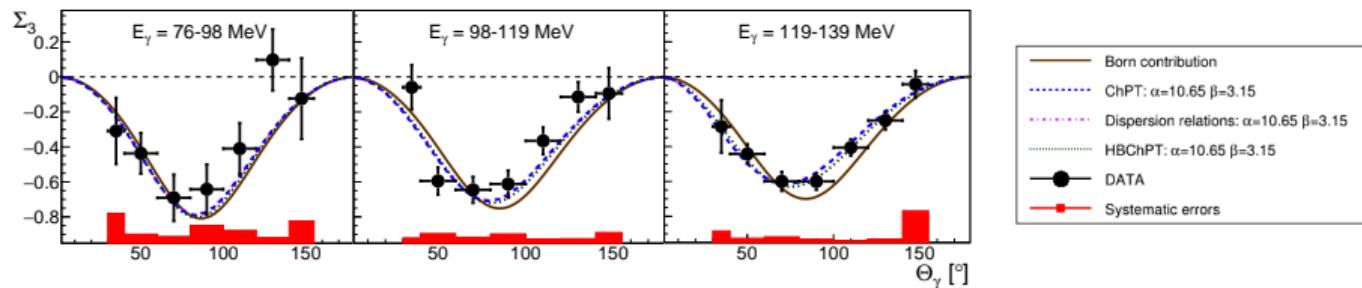


# Pilot experiment

V. Sokhyan, E.J. Downie, E. Mornacchi, J.A. McGovern, N. Krupina et al., Eur. Phys. J. A (2017) 53:14

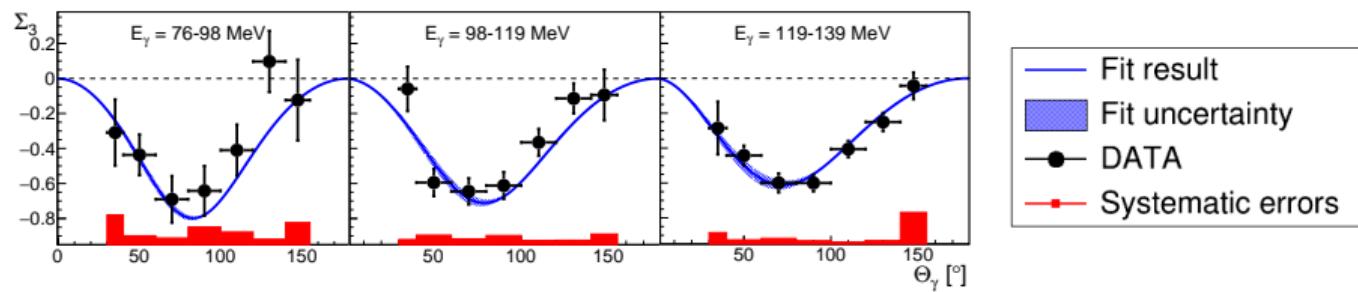
~200k good Compton scattering event in the range  $E_{\gamma b} = 80 - 140$  MeV

- Theoretical predictions for fixed  $\alpha_{E1}$  and  $\beta_{M1}$ :



Lensky, V. & Pascalutsa, V.,  
Eur. Phys. J. C (2010) 65:195  
McGovern, J.A., Phillips, D.R. &  
Grießhammer, H.W.,  
Eur. Phys. J. A (2013) 49:12  
B. Pasquini, D. Drechsel, and M.  
Vanderhaeghen, Phys. Rev. C 76

- Fit results using only new  $\Sigma_3$  data within ChPT framework:



Lensky, V. & Pascalutsa, V.,  
Eur. Phys. J. C (2010) 65:195

## Low-energy expansion

The low-energy expansion, developed by Petrun'kin, calculate the Compton scattering amplitude to the order of  $\omega^2$ :

$$\left( \frac{d\sigma}{d\Omega} \right) = \left( \frac{d\sigma}{d\Omega} \right)_{\text{Born}} + \left( \frac{d\sigma}{d\Omega} \right)_{\text{NB}},$$

where:

$$\left( \frac{d\sigma}{d\Omega} \right)_{\text{Born}} = \frac{1}{2} \left( \frac{e^2}{m} \right)^2 \left( \frac{\omega'}{\omega} \right)^2 \times \left\{ 1 + \cos^2 \theta_{\gamma'} + \frac{\omega \omega'}{m^2} ([1 - \cos \theta_{\gamma'}]^2 + a_0 + a_1 \cos^2 \theta_{\gamma'} + a_2 z^2) \right\},$$

$$\left( \frac{d\sigma}{d\Omega} \right)_{\text{NB}} = -\omega \omega' \left( \frac{\omega'}{\omega} \right)^2 \frac{e^2}{m} \left[ \frac{\alpha_{E_1} + \beta_{M_1}}{2} (1 + \cos \theta_{\gamma'})^2 + \frac{\alpha_{E_1} + \beta_{M_1}}{2} (1 - \cos \theta_{\gamma'})^2 \right].$$

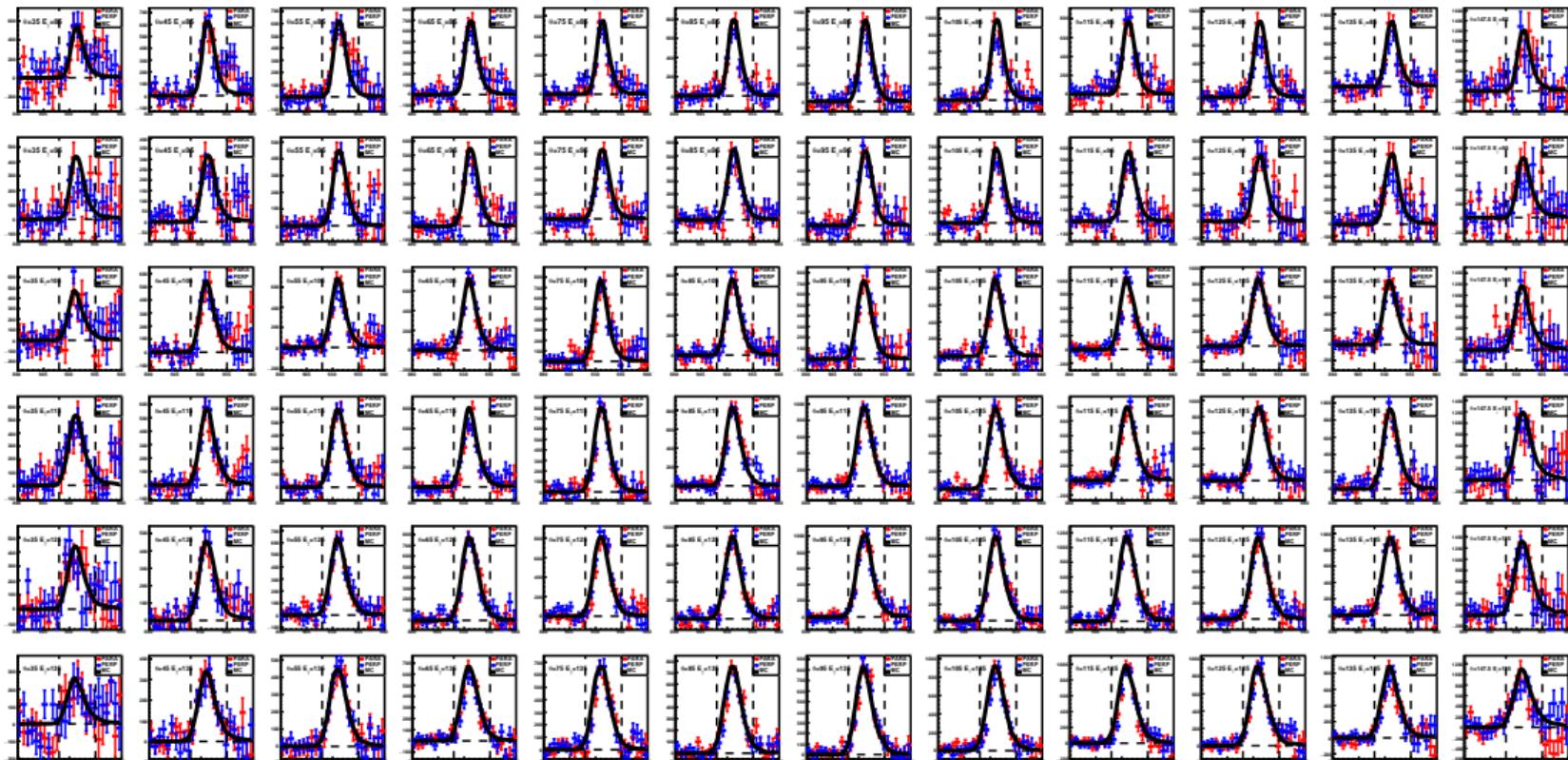
Here  $\omega'$  is the energy of the scattered photon, given by:

$$\omega' = \frac{\omega}{1 + (\omega/m)(1 + \cos \theta_{\gamma'})},$$

while the coefficients  $a_0$ ,  $a_1$  and  $a_2$  are combination of the anomalous magnetic moment  $k$ .

# Missing Mass empty target subtracted

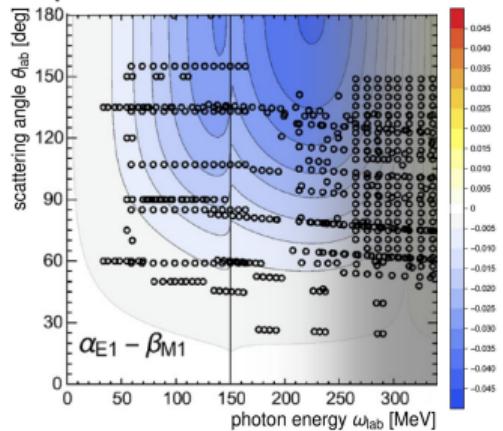
A2



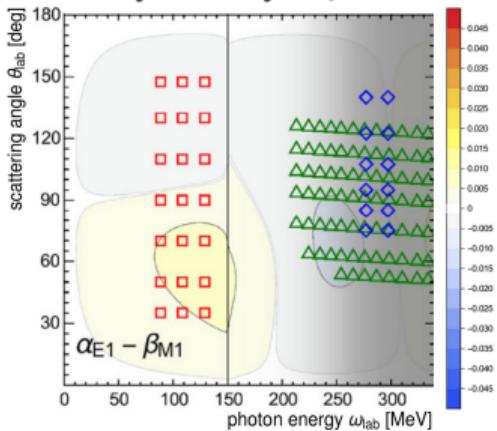
# Sensitivity

A2

Unpolarized cross-section



Beam asymmetry  $\Sigma_3$



Grießhammer, H.W., McGovern, J.A. & Phillips, D.R., Eur. Phys. J. A (2018) 54:37

# Motivation - muonic hydrogen 2S-2P Lamb shift

A2

The two-photon exchange contributions to the muonic hydrogen Lamb shift are usually divided into an elastic part and an inelastic part (proton polarization contribution  $\Delta E_P$ ).

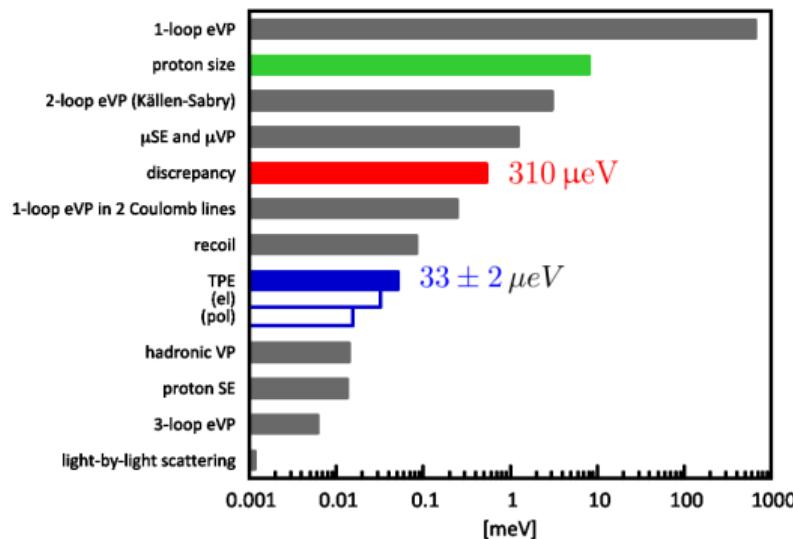


TABLE I. Numerical results for the  $O(\alpha^5)$  proton structure corrections to the Lamb shift in muonic hydrogen. Energies are in  $\mu$ eV.

( $\mu$ eV)	This work	Refs. [11,12]	Ref. [22]
$\Delta E_{\text{subt}}$	$5.3 \pm 1.9$	1.8	2.3
$\Delta E_{\text{incl}}$	$-12.7 \pm 0.5$	$-13.9$	$-13.8$
$\Delta E_{\text{cl}}$	$-29.5 \pm 1.3$	$-23.0$	$-23.0$
$\Delta E$	$-36.9 \pm 2.4$	$-35.1$	$-34.5$

C. E. Carlson & M. Vanderhaeghen, Phys. Rev. A 84 (2011)

[11] K. Pachucki, Phys. Rev. A 53 (1996)

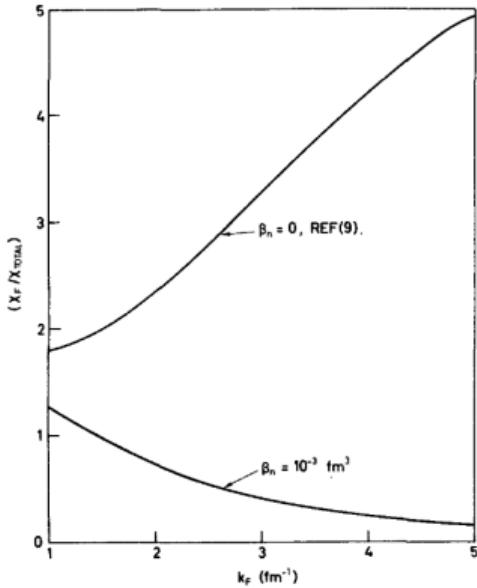
[12] K. Pachucki, Phys. Rev. A 53 (1999)

[22] A. P. Martynenko, Phys. At. Nucl. 69 (2006)

F. Hagelstein, R. Miskimen & V. Pascalutsa, Progress in Particle and Nuclear Physics 88 (2016)

## Motivation - neutron star

In the case of  $\beta_{M1}^N \approx 10^{-3}$  fm<sup>3</sup>, the intrinsic neutron polarizability  $\chi_n = \beta_n \rho = \beta_n k_F^3 / 3\pi^2$  is the dominant contribution to the neutron star susceptibility, compare to the Pauli paramagnetic contribution  $\chi_F = \mu^2 m_n k_F / \pi$



Effect of the magnetic polarizability on the paramagnetic susceptibility as a function of the Fermi wave vector

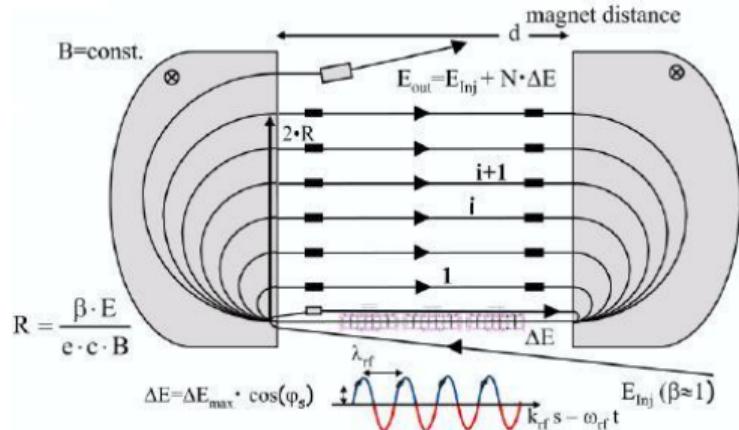
J. Bernabeu, T. E. O. Ericson, and C. Ferro Fontan, Phys. Lett. B, 49:381, 1974.

The determination of the electromagnetic self-energy contribution to the proton-neutron mass difference is limited by our knowledge of the difference between the proton and neutron scalar magnetic polarizabilities, for which even the sign is presently unknown.

A. Walker-Loud, C. E. Carlson, and G. A. Miller, Phys. Rev. Lett. 108 (2012).

# Race Track Microtron

A2



- Every lap, the electrons are accelerated by the LINAC:

$$\Delta E = \frac{ec^2 B}{2\pi\nu_{fr}}$$

- The final energy is give by:

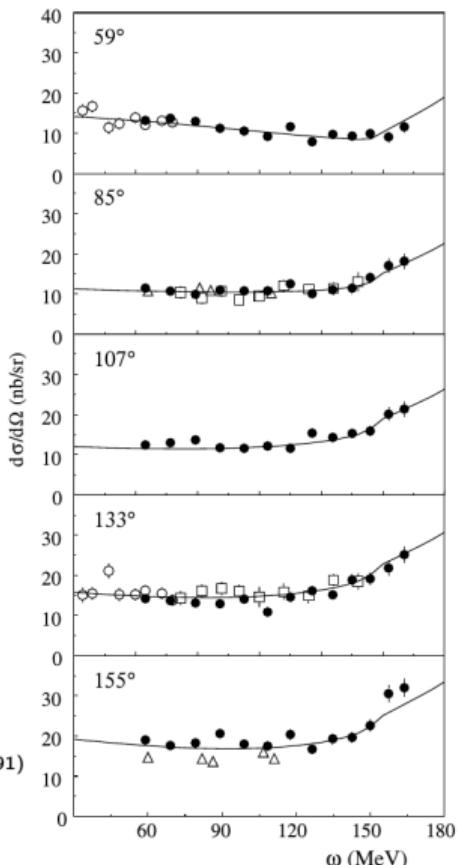
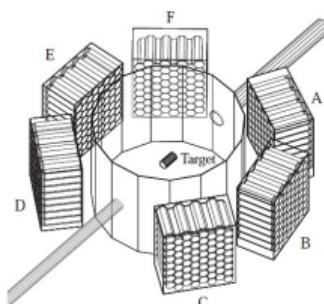
$$\Delta E_{\text{out}} = E_i + N\Delta E$$

# Existing data on Compton scattering on the proton - TAPS dataset

A2

- Highest statistics published data:  
V. Olmos de Leon et al. Eur. Phys. J.  
A 10, 207-215 (2001)
- 200 hours of Compton scattering
- $E_{\text{beam}} = 180 \text{ MeV}$
- $E_{\gamma} = 55 - 165 \text{ MeV}, \theta_{\gamma} = 59^\circ - 155^\circ$
- $\sim 1/3$  acceptance of the A2 apparatus

$$\alpha_{E1} = (12.1 \pm 1.08) 10^{-4} \text{ fm}^3$$
$$\beta_{M1} = (1.6 \mp 0.89) 10^{-4} \text{ fm}^3$$

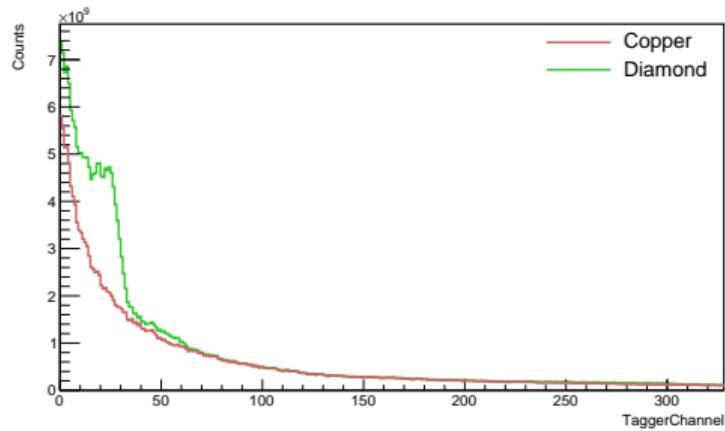


Triangles: P.S. Baranov et al.(1974)  
P.S. Baranov et al.(1975)  
Open circles: F.J. Federspiel et al.(1991)  
Squares B.E. MacGibbon et al.(1995)  
Curve: R.A. Arndt et al.(1996)

# Polarized photon beam

A high energy electron can produce Bremsstrahlung photons when slowed down by a dense material.

- From an unpolarized electron beam, one can produce **unpolarized** photons using an amorphous radiator or **polarized** photons using a diamond radiator

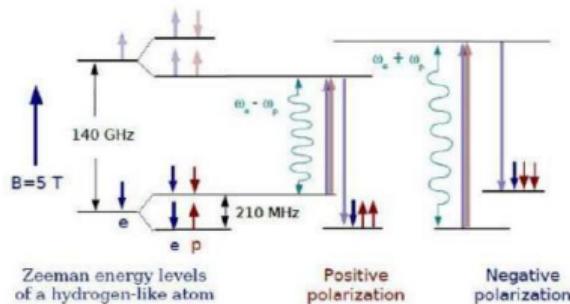


- From a longitudinally polarized electron beam, one can produce circularly polarized photons

# Frozen spin target

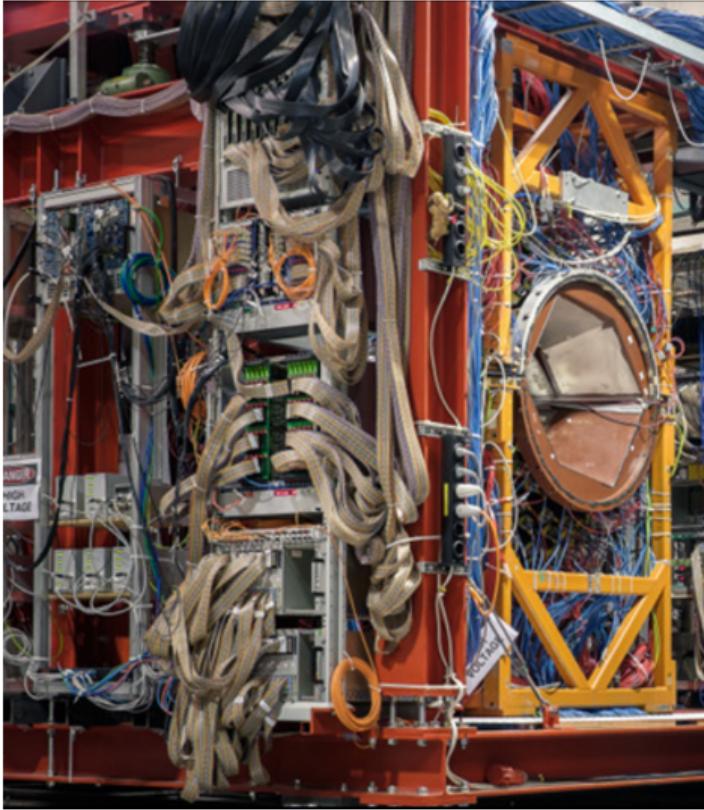
Protons are polarized via Dynamic Nuclear Polarization (DNP):

- cool target to 0.2 K
- use 2.5 T to align electron spins
- pump  $\approx 70$  GHz microwaves (just above, or below, the electron spin resonance frequency), causing spin-flips between electrons and protons
- cool target to 0.025 K in order to freeze the proton spins
- remove the polarizing magnet and turn on 0.6 T in the cryostat to maintain the polarization
- get a relaxation time  $> 1000$  h and a polarization up to 90%



# Detectors - Crystal Ball system

A2

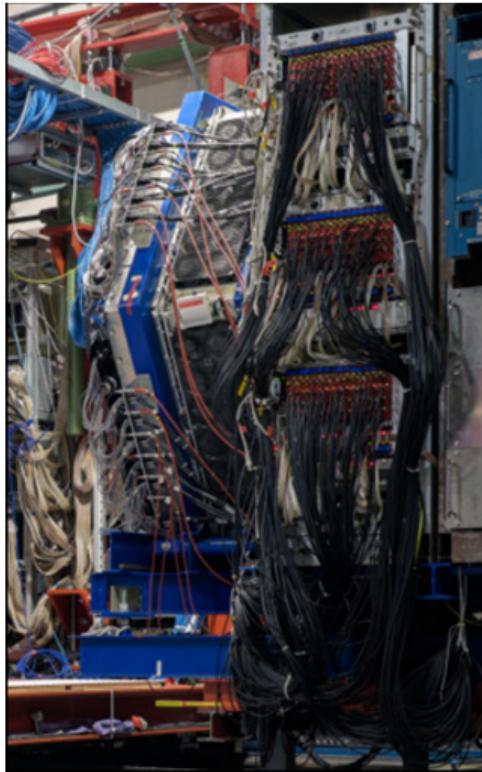


- Proposed at SLAC in 1974
- Used in SLAC, DESY, Brookhaven
- MAMI since 2002
- 672 NaI detectors
- $21^\circ < \theta < 159^\circ$
- Full  $\phi$  coverage
- E resolution  $\approx 3\%$
- $\theta$  resolution  $\approx 2.5^\circ$

# Detectors - TAPS

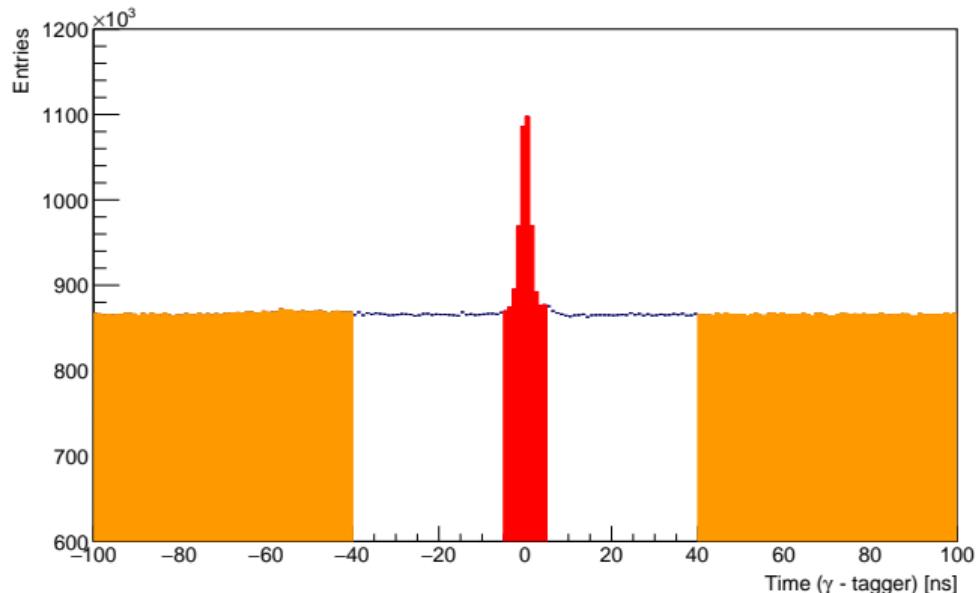
A2

- Built in 1980s from TAPS collaboration
- Designed from many experiments in different configurations
- 366 BaF<sub>2</sub> detectors
- 72 PbWO<sub>4</sub> detectors
- Covers  $\theta < 20^\circ$
- E resolution  $\approx 3\%$
- $\theta$  resolution  $\approx 0.7^\circ$



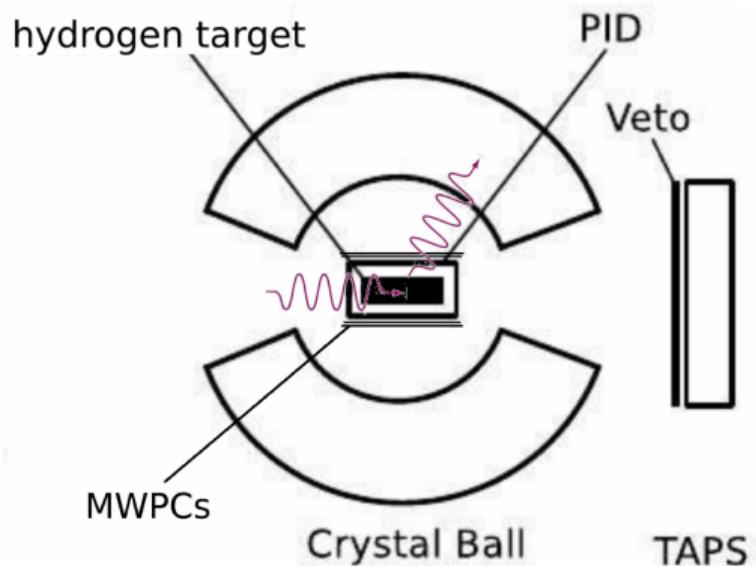
# Random coincidences subtraction

Cut on **prompt** and **random** photons in the time spectra. **Randoms** are scaled according to the time interval and subtracted from the **prompts**, in order to remove accidental coincidences.



## Event selection

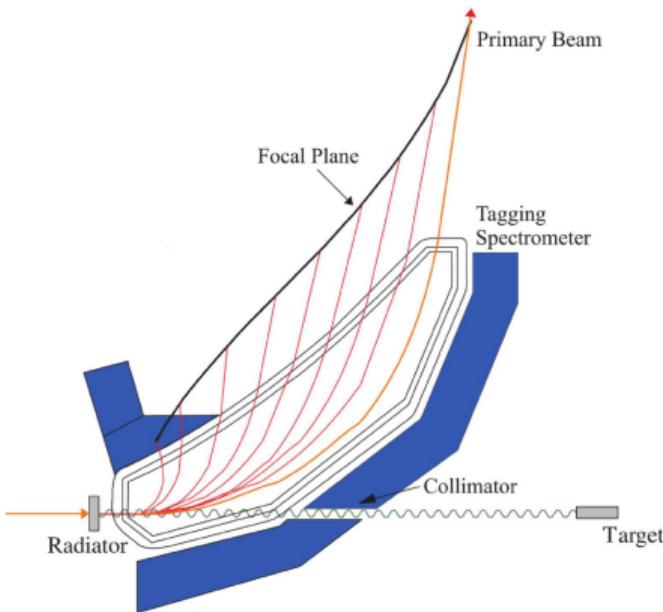
Select event with **ONLY one photon**. At this energy, the proton stops inside the target!



# Photon tagging system

High intensity beam of linearly polarized tagged photons:

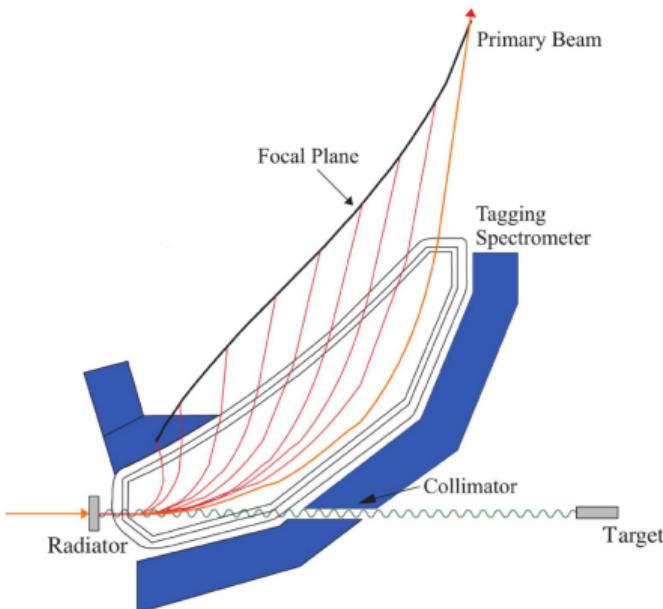
$$E_\gamma = E_0 - E_{e^-}$$



# Photon tagging system

High intensity beam of linearly polarized tagged photons:

$$E_\gamma = E_0 - E_{e^-}$$



## Upgrade of the focal plane detector

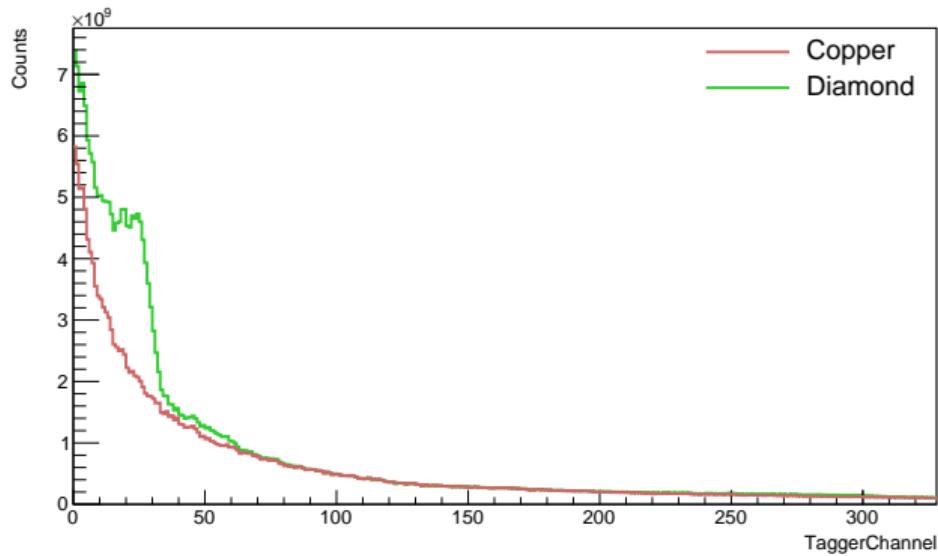
- Higher photon flux
- Higher efficiency
- Better control of systematic



# Linear polarization

A2

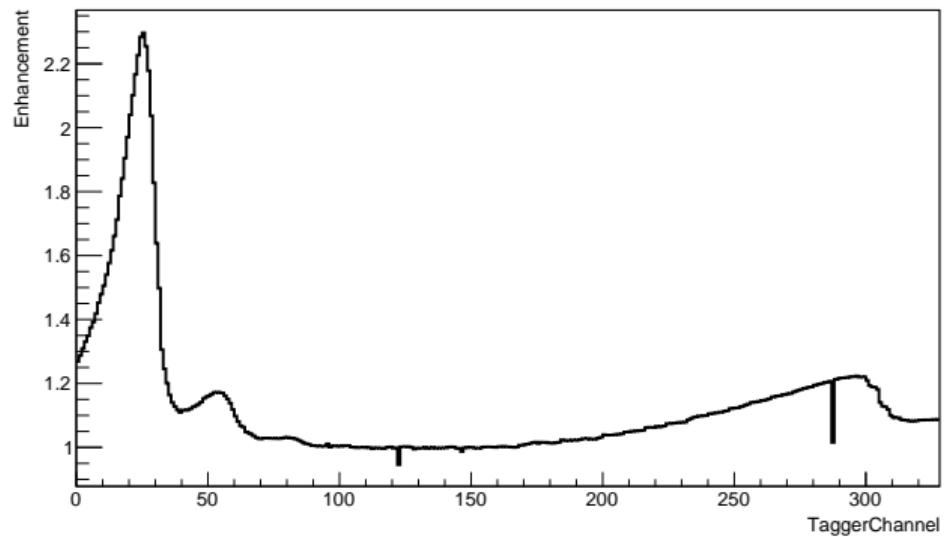
The **amorphous** reference is taken every  $\sim 24$  hours to account for possible variation in the bremsstrahlung distribution



# Linear polarization

A2

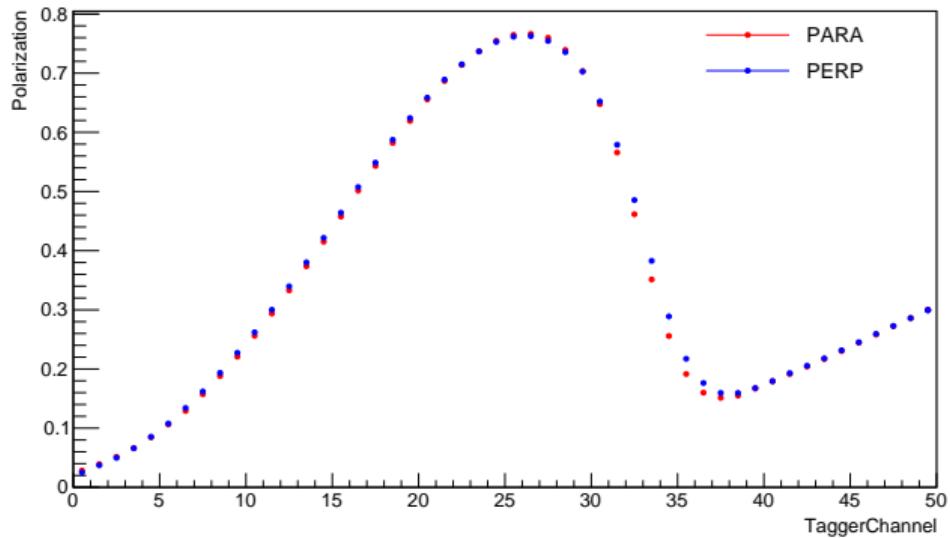
The enhancement is obtained dividing the **polarized** and the **amorphous**. It is then fitted every  $\sim 2$  seconds to determine the polarization



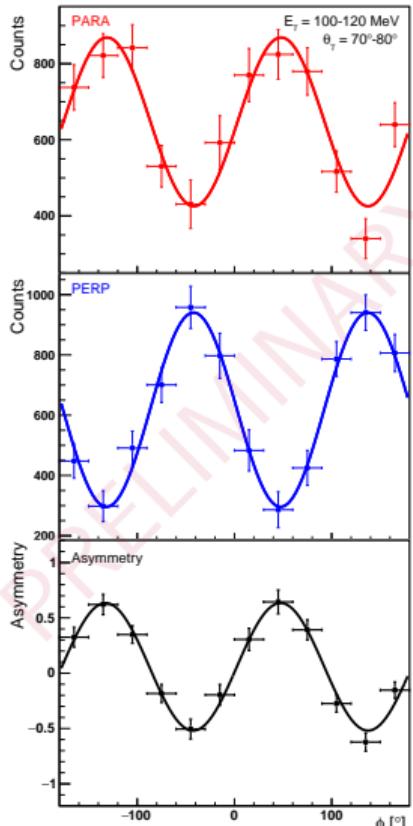
# Linear polarization

A2

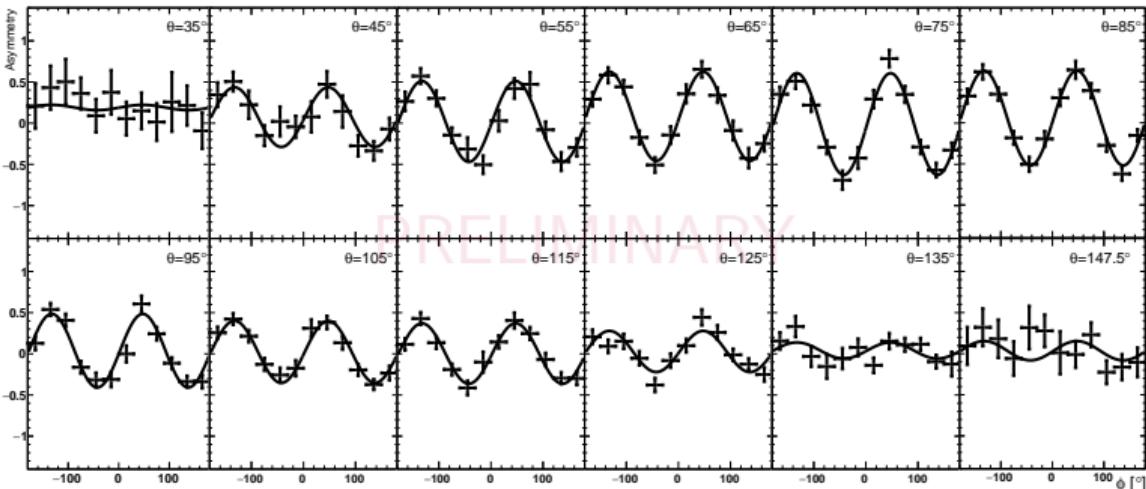
The polarization degree is then determined for different beam energies both for the **parallel** and the **perpendicular** dataset



# New measurement - $\phi$ distributions



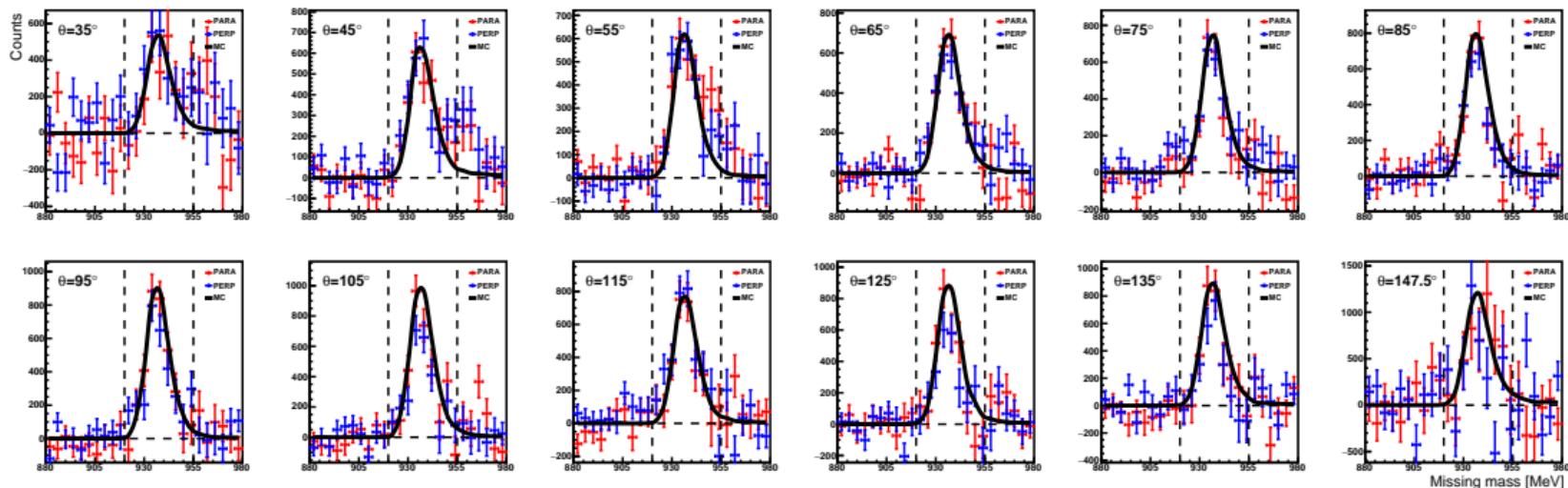
$E_\gamma = 130 \text{ MeV}$ :



$$\frac{d\sigma}{d\Omega}(E, \theta, \phi) = \frac{d\sigma}{d\Omega}(E, \theta) [1 + p_\gamma(E) \Sigma_3(E, \theta) \cos(2\phi)]$$

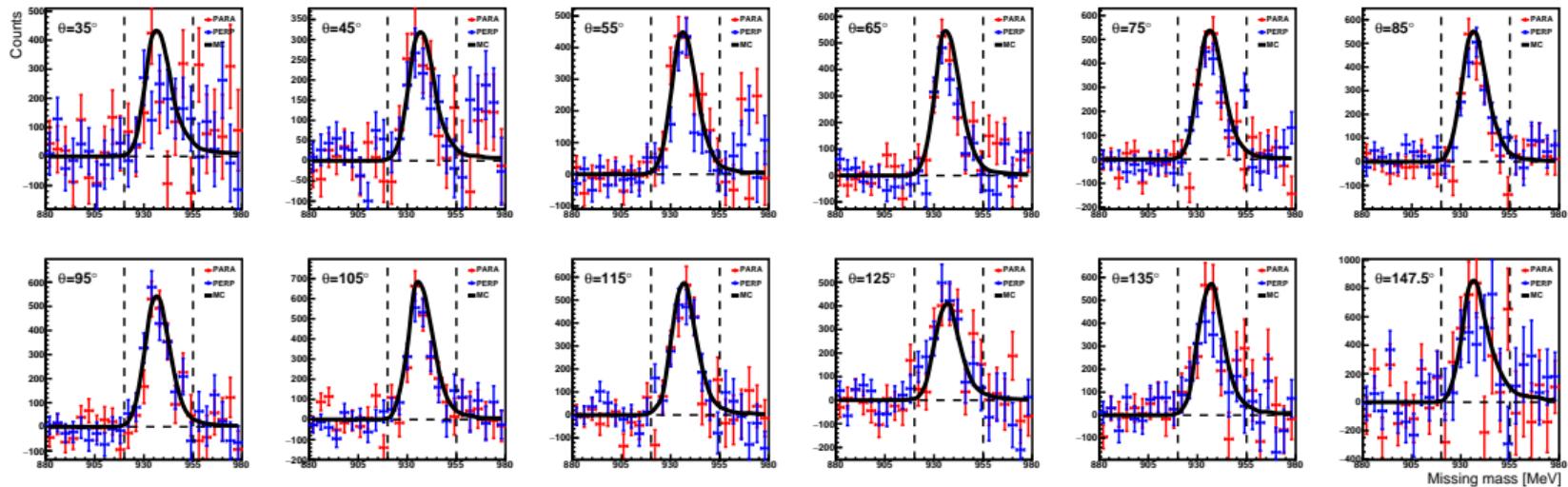
# Missing mass $E_\gamma = 85$ MeV

A2



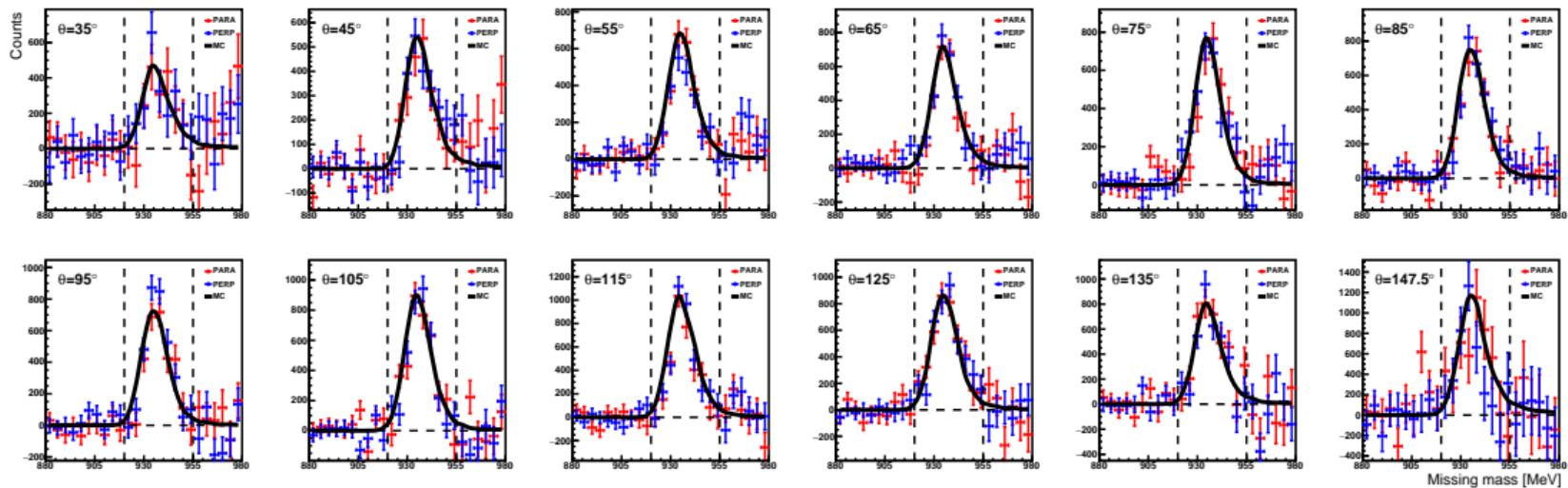
# Missing mass $E_\gamma = 95$ MeV

A2



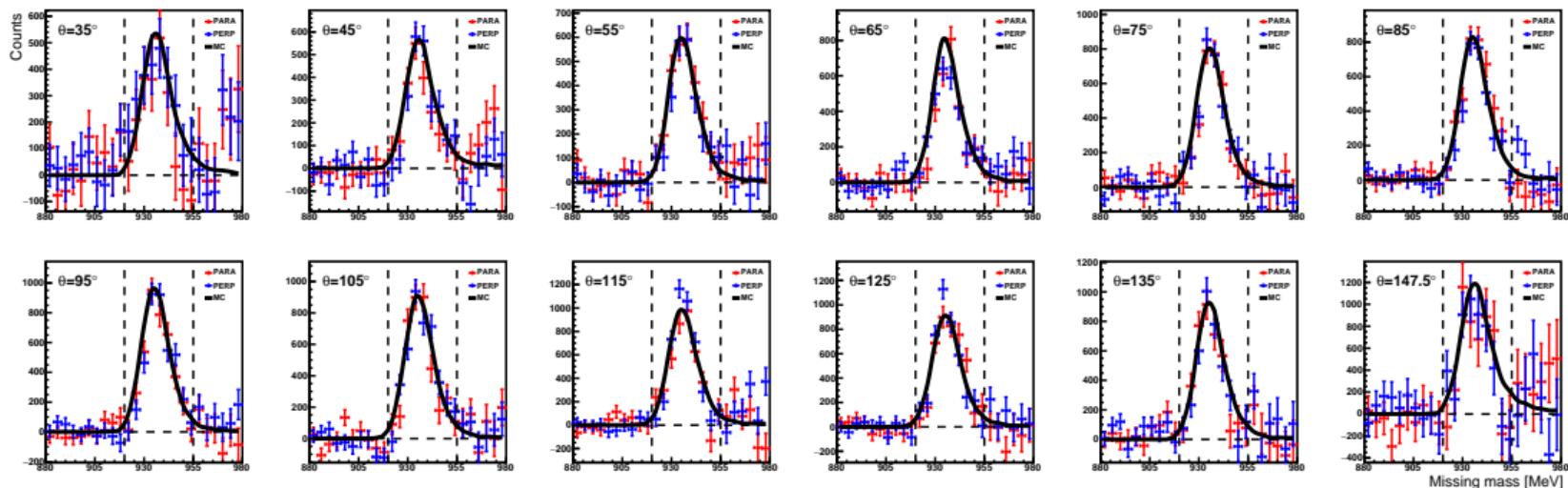
# Missing mass $E_\gamma = 105$ MeV

A2



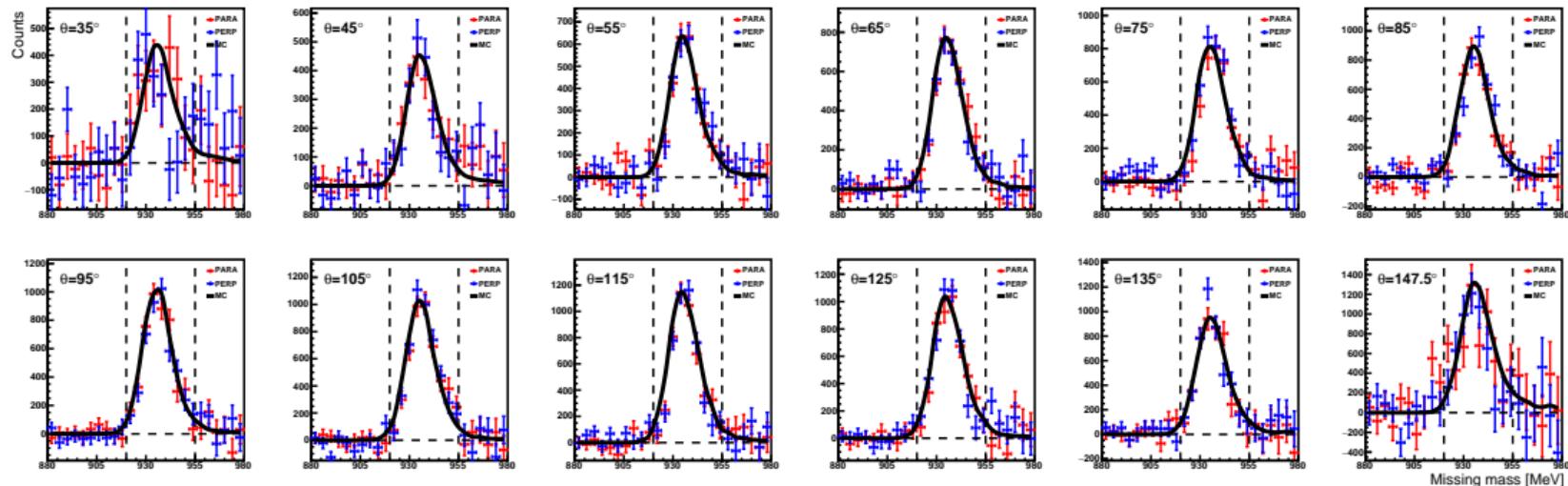
# Missing mass $E_\gamma = 115$ MeV

A2



# Missing mass $E_\gamma = 125$ MeV

A2



# Missing mass $E_\gamma = 135$ MeV

A2

